

Bose-Einstein condensate in an elliptical waveguide

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Abstract

We investigate the effects of spatial curvature for an atomic Bose-Einstein condensate confined in an elliptical waveguide. The system is well described by an effective 1D Gross-Pitaevskii equation with a quantum-curvature potential, which has the shape of a double-well but crucially depends on the eccentricity of the ellipse. The ground state of the system displays a quantum phase transition from a two-peak configuration to a one-peak configuration at a critical attractive interaction strength. In correspondence of this phase transition the superfluid fraction strongly reduces and goes to zero for a sufficiently attractive Bose-Bose interaction.



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Published by the SciPost Foundation.

Received 18-09-2021

Accepted 18-03-2022

Published 22-03-2022

doi:[10.21468/SciPostPhysCore.5.1.015](https://doi.org/10.21468/SciPostPhysCore.5.1.015)



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1 Introduction

How does a locally-varying spatial curvature influence the properties of low-dimensional quantum systems? This is a relevant question asked by scientists working in very different fields such as the linear Schrödinger equation for a particle constrained on a curve manifold [1–3], but also quantum gravity [4] or quantum chaos [5]. It is well known [6–8] that the local curvature of a curve on the three-dimensional (3D) Euclidean space is characterized by the so-called geodesic curvature. This geodesic curvature κ is an extrinsic quantity: it does not remain invariant if the curve is under the effect of a distance-preserving transformation [6–8]. Instead, the local curvature of a surface on the 3D Euclidean space is characterized by the so-called Riemann curvature tensor, which can be written in terms of the invariant Gaussian curvature and the not-invariant average curvature [6–8]. The quantum motion of a particle on a curved waveguide has been analyzed by several authors [9–14]. More recently, the highly non-trivial role of curvature for constrained quantum systems has been theoretically investigated with ultracold atomic gases confined in a quasi-1D [15–18] and quasi-2D configurations [19]. The main result of all these investigations is that the local curvature gives rise to a quantum-curvature potential.

In this paper we consider an atomic Bose-Einstein condensate (BEC) confined in a quasi-1D elliptical waveguide finding that the quantum-curvature potential has the shape of a double-well, if the eccentricity of the ellipse is different from zero. By numerically solving the 1D Gross-Pitaevskii equation of the BEC wavefunction under the effect of this quantum-curvature potential, we show that the ground state of the system is uniform along the waveguide only if the eccentricity ϵ of the ellipse is zero (circular waveguide with constant curvature). Instead, for $\epsilon \neq 0$ we find that the ground state is generically characterized by a two-peak configuration, where the peaks are located around the minima of the effective double-well potential. However, we discover that in the case of attractive interaction it exists a critical (negative) interaction strength below which the ground state exhibits a quantum phase transition from the two-peak configuration to a one-peak configuration. This is the analog of the spontaneous symmetry breaking, i.e. the modulational instability [20], of the uniform configuration predicted some years ago for an 1D attractive BEC in a circular waveguide [21–23]. Our results show that the critical interaction strength depends on the eccentricity ϵ of the ellipse in a non-trivial way. We also analyze the effect of a boost velocity on the BEC moving in the elliptical waveguide deriving the Leggett formula [24] for the superfluid fraction of a 1D bosonic system [25–27]. Our numerical investigation reveals that the superfluid fraction decreases dramatically in response to this quantum phase transition, eventually reaching zero for a sufficiently negative Bose-Bose interaction.

2 Quantum-curvature potential

We consider a Bose-Einstein condensate (BEC) made of N identical bosonic atoms of mass m . The atoms are constrained to move along a curve \mathcal{C} by the presence of a strong harmonic potential of frequency ω_{\perp} in the local transverse plane with respect to \mathcal{C} . The characteristic length of the transverse confinement is $l_{\perp} = \sqrt{\hbar/(m\omega_{\perp})}$ where \hbar is the reduced Planck constant. We introduce a local system (s, u, v) of coordinates, where s is the curvilinear abscissa (arclength) along \mathcal{C} while u and v are two coordinates of the transverse plane [15–18]. In this way the Lagrangian density of our problem is given by

$$\mathcal{L} = \frac{i\hbar}{2} (\Psi^* \partial_t \Psi - \Psi \partial_t \Psi^*) - \frac{\hbar^2}{2m} |\nabla \Psi|^2 - \frac{m\omega_{\perp}^2}{2} (u^2 + v^2) |\Psi|^2 - \frac{1}{2} g |\Psi|^4, \quad (1)$$

where $\Psi(s, u, v, t)$ is the BEC wavefunction normalized to one and $g = 4\pi\hbar^2 a_s(N-1)/m$ is the 3D strength of the contact inter-atomic potential with a_s the s-wave scattering length. Clearly, the Laplacian operator ∇^2 must be written in terms of the local system (s, u, v) of coordinates [15–18]. Assuming the factorization

$$\Psi(s, u, v, t) = \psi(s, t) \frac{e^{-\frac{(u^2+v^2)}{2\sigma(s,t)^2}}}{\pi^{1/2}\sigma(s, t)} \quad (2)$$

and inserting this ansatz into the Lagrangian density, after integration over u and v one gets [18, 28, 29],

$$\begin{aligned} \bar{\mathcal{L}} = & \frac{i\hbar}{2} (\psi^* \partial_t \psi - \psi \partial_t \psi^*) - \frac{\hbar^2}{2m} |\partial_s \psi|^2 + \frac{\hbar^2 \kappa^2(s)}{8m} |\psi|^2 \\ & - \left(\frac{\hbar^2}{2m} \frac{1}{\sigma^2} + \frac{m\omega_\perp^2}{2} \sigma^2 \right) |\psi|^2 - \frac{1}{2} \frac{g}{2\pi\sigma^2} |\psi|^4, \end{aligned} \quad (3)$$

where $\kappa(s)$ is the local geodesic curvature of \mathcal{C} , and the conditions $\sigma\kappa \ll 1$ and $\sigma \ll \xi$ must hold, with $\xi = \hbar/\sqrt{2g|\Psi|^2}$ the 3D healing length [18]. The Euler-Lagrange equations of the 1D action functional with respect to the 1D wavefunction $\psi(s, t)$ and the transverse width $\sigma(s, t)$ are

$$i\hbar\partial_t \psi = \left[-\frac{\hbar^2}{2m} \partial_s^2 - \frac{\hbar^2 \kappa^2(s)}{8m} + \frac{\hbar^2}{2m} \frac{1}{\sigma^2} + \frac{m\omega_\perp^2}{2} \sigma^2 + \frac{2\hbar^2 a_s(N-1)}{m\sigma^2} |\psi|^2 \right] \psi, \quad (4)$$

and

$$\sigma^2 = l_\perp^2 \sqrt{1 + 2a_s(N-1)|\psi|^2}. \quad (5)$$

Eq. (4), equipped with Eq. (5), is the time-dependent 1D nonpolynomial Schrödinger equation (NPSE) [28, 29] for the wavefunction $\psi(s, t)$ of the BEC moving along the curve \mathcal{C} (see also [18]). As previously discussed, the geodesic curvature $\kappa(s)$ gives rise to an effective potential

$$U_Q(s) = -\frac{\hbar^2 \kappa(s)^2}{8m}. \quad (6)$$

This curvature potential $U_Q(s)$ is quantum because it involves the square of the reduced Planck constant \hbar . At fixed atomic mass m , only if the square of the curvature $\kappa(s)$ is sufficiently large the effects of this quantum-curvature potential become relevant.

Under the assumption that $\sigma \simeq l_\perp$, which corresponds to a very strong transverse confinement, the 1D NPSE becomes the familiar 1D Gross-Pitaevskii (GPE) equation

$$i\hbar\partial_t \psi = \left[-\frac{\hbar^2}{2m} \partial_s^2 - \frac{\hbar^2 \kappa(s)^2}{8m} + \hbar\omega_\perp + \frac{2\hbar^2 a_s(N-1)}{ml_\perp^2} |\psi|^2 \right] \psi. \quad (7)$$

It is very important to stress that, from Eq. (5), the condition $\sigma \simeq l_\perp$ implies $2a_s(N-1)|\psi|^2 \ll 1$. In the rest of the paper we will work within this 1D regime. In the new version of the manuscript I shall discuss the role. Clearly, Eq. (7) is reliable in the weak-coupling and strong-transverse-confinement regime, where both beyond-mean-field and transverse-size effects are very small.

3 Properties of the elliptical waveguide

We now choose an ellipse for the curve \mathcal{C} . By using cartesian coordinates its defining equation reads

$$\frac{x^2}{a^2} + \frac{y^2}{b^2} = 1, \quad (8)$$

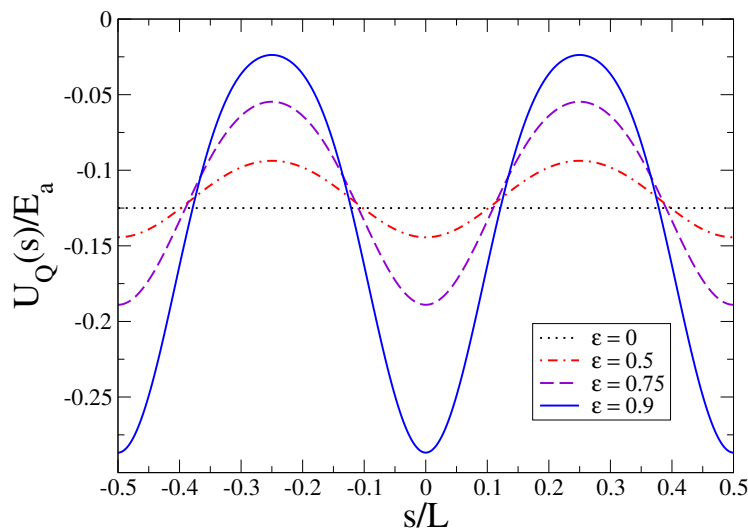


Figure 1: Quantum-curvature potential $U_Q(s)$, Eq. (6), induced by the geodesic curvature $\kappa(s)$ of an ellipse, as a function of the arclength s , where a is the length of the major semi-axis, $L = aE(2\pi, \epsilon)$ is the perimeter of the ellipse, and $E_a = \hbar^2/(ma^2)$ a characteristic energy. The curves are obtained for different values of the eccentricity ϵ .

where a and b are the lengths of the two semi-axes of the ellipse. Here we assume that $a \geq b$, such that a is the length of the major semi-axis. The eccentricity of the ellipse is defined as [8]

$$\epsilon = \sqrt{1 - \frac{b^2}{a^2}}. \quad (9)$$

Clearly, $0 \leq \epsilon < 1$ and for $\epsilon = 0$ we obtain a circle of radius $R = a = b$. Introducing the angle $\phi \in [0, 2\pi]$ we can write

$$x = a \cos(\phi), \quad (10)$$

$$y = b \sin(\phi), \quad (11)$$

and the arclength s along the ellipse can be expressed with the formula [8]

$$s = a E(\phi, \epsilon), \quad (12)$$

where

$$E(\phi, \epsilon) = \int_0^\phi \sqrt{1 - \epsilon^2 \sin^2(\phi')} d\phi' \quad (13)$$

is the incomplete elliptic integral of the second kind. It follows that the perimeter L of the ellipse reads

$$L = a E(2\pi, \epsilon), \quad (14)$$

such that for $\epsilon = 0$ we have $L = 2\pi a$ because $E(2\pi, 0) = 2\pi$. Instead, for $\epsilon \rightarrow 1$ we have $L \rightarrow 4a$ because $E(2\pi, 1) = 4$. We conclude that $4a < L \leq 2\pi a$. The geodesic curvature κ of the ellipse can be written as [8]

$$\kappa = \frac{1}{a} \frac{\sqrt{1 - \epsilon^2}}{(\sin^2(\phi) + \sqrt{1 - \epsilon^2} \cos^2(\phi))^{3/2}}. \quad (15)$$

At fixed a and ϵ , the maximum of the curvature is obtained for $\phi = 0$ and $\phi = \pi$, i.e.

$$\kappa_{max} = \frac{1}{a} \frac{1}{(1 - \epsilon^2)^{1/4}}, \quad (16)$$

while the minimum of the curvature is obtained for $\phi = \pi/2$ and $\phi = 3\pi/2$, i.e.

$$\kappa_{min} = \frac{1}{a} \sqrt{1 - \epsilon^2}. \quad (17)$$

Thus, for $\epsilon = 0$ we have $\kappa_{max} = \kappa_{min} = 1/a$ while for $\epsilon \rightarrow 1$ we have $\kappa_{max} \rightarrow +\infty$ and $\kappa_{min} \rightarrow 0$. We conclude that $1/a \leq \kappa_{max} < +\infty$ and $0 < \kappa_{min} \leq 1/a$. In general, the formula which gives the curvature κ as a function of the arclength s is called Cesaro equation. Unfortunately, in the case of the ellipse there is no Cesaro equation. In other words, an analytical formula of κ as a function of s is not available. However, from Eqs. (12) and (15), fixing the length a and the eccentricity ϵ of the ellipse, we can easily plot κ vs s using ϕ as dummy variable. More explicitly: we calculate separately κ vs ϕ and s vs ϕ , and then we plot κ vs s . The curvature $\kappa(s)$ has a the periodic structure of $\kappa(s)$. By increasing ϵ , the perimeter L of the ellipse slightly decreases while κ_{max} and κ_{min} pull away. In Fig. 1 we plot the quantum-curvature potential $U_Q(s)$, Eq. (6), induced by the curvature $\kappa(s)$ of the ellipse for different values of the eccentricity ϵ . The figure clearly shows that, for $\epsilon \neq 0$, $U_Q(s)$ is symmetric double-well potential where the depth of the wells becomes larger by increasing the eccentricity. The minima (maxima) of the quantum-curvature potential $U_Q(s)$ are in correspondence to the maxima (minima) of the curvature $\kappa(s)$.

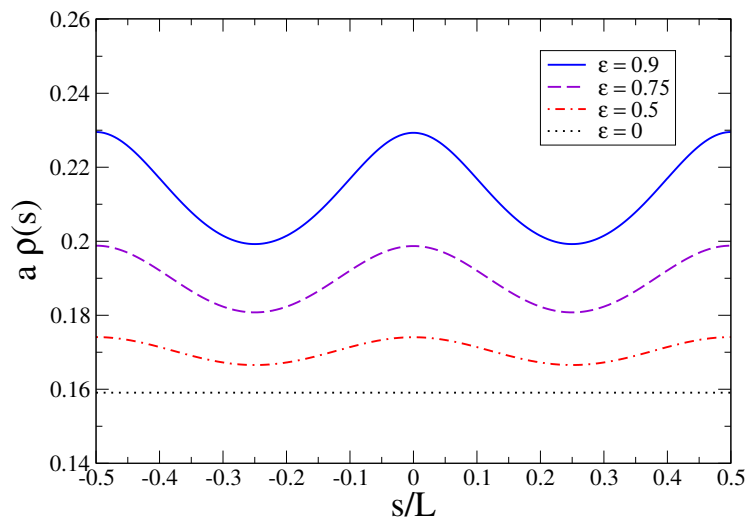


Figure 2: Probability density $\rho(s)$ of the non-interacting BEC ground state in an ellipse as a function of the arclength s , where a is the length of the major semi-axis and $L = aE(2\pi, \epsilon)$ is the perimeter of the ellipse. The curves are obtained for different values of the eccentricity ϵ . Here the s -wave scattering length a_s is set to zero or, equivalently, the number N of particles is set to one.

4 BEC ground state in elliptical waveguide

The time-independent 1D GPE is obtained from Eqs. (7) setting

$$\psi(s, t) = \Phi(s) e^{-i(\mu + \hbar\omega_{\perp})t/\hbar}. \quad (18)$$

In this way we have

$$\mu \Phi = \left[-\frac{\hbar^2}{2m} \partial_s^2 - \frac{\hbar^2 \kappa(s)^2}{8m} + \frac{2a_s(N-1)\hbar^2}{ml_\perp^2} |\Phi|^2 \right] \Phi, \quad (19)$$

that is the 1D GPE equation for the stationary wavefunction $\Phi(s)$, such that $\rho(s) = |\Phi(s)|^2$ is the probability density of finding the BEC at the position s . In Fig. 2 we report the probability density $\rho(s)$ of the ground-state of the non-interacting ($a_s = 0$ or, equivalently, $N = 1$) BEC confined along the ellipse as a function of the arclength s . Our results are obtained by solving Eq. (7) with a Crank-Nicolson predictor-corrector method and imaginary time. In the figure, the curves correspond to different values of the eccentricity ϵ . Clearly, for $\epsilon = 0$ the ground state is uniform along the ellipse. However, for $\epsilon \neq 0$ the ground state is no more uniform due to a non-constant curvature $\kappa(s)$ which implies a non-constant effective potential $U_Q(s) = -\hbar^2 \kappa(s)^2 / (8m)$. By increasing the eccentricity ϵ the localization of $\rho(s)$ around the minima of $U_Q(s)$, where the curvature is larger, becomes more evident.

5 Quantum phase transition

It is interesting to investigate the effect of the inter-atomic interaction on the ground state properties of the system. In adimensional units the interaction strength reads $\gamma = g / (2\pi l_\perp^2 a E_a) = 2aa_s(N-1)/l_\perp^2$ with $E_a = \hbar^2 / (ma^2)$. In Fig. 3 we consider the case of an attractive BEC and plot our numerical results obtained for different values of a negative γ with fixed eccentricity $\epsilon = 0.9$. Quite remarkably, for $\gamma < -1.5$ the ground state has a spontaneous symmetry breaking: one of the two local minima contains more bosons. Indeed for $\gamma = -2$ this single-well localization is very clear. It is important to observe that this kind of quantum phase transition happens for any ϵ . In the case of a circle ($\epsilon = 0$) a similar spontaneous symmetry breaking was predicted about 20 years ago [21–23]. Actually, this quantum

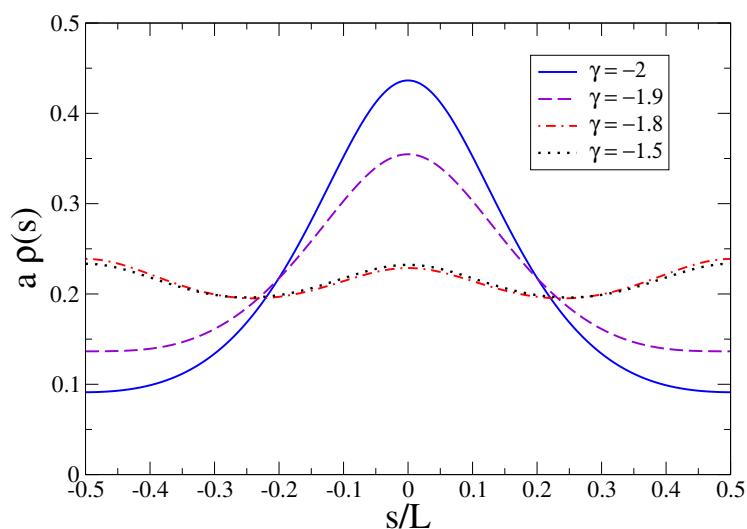


Figure 3: Probability density $\rho(s)$ of the attractive BEC ground state in an ellipse as a function of the arclength s , where a is the length of the major semi-axis and $L = aE(2\pi, \epsilon)$ is the perimeter of the ellipse. The curves are obtained with eccentricity $\epsilon = 0.9$ for different values of the adimensional interaction strength $\gamma = 2aa_s(N-1)/l_\perp^2$, where a_s is the 3D s-wave scattering length and l_\perp is the characteristic length of the transverse harmonic confinement.

phase transition, or spontaneous symmetry breaking, is nothing else than the modulational instability of the ground-state configuration, induced by the appearance of an imaginary component in the energies of the elementary excitations of the ground state [20]. However, for $\epsilon = 0$ there is a quantum phase transition from a uniform configuration to a single-peak configuration, while for the $\epsilon \neq 0$ there is a quantum phase transition from a two-peak configuration to a single-peak configuration. This quantum phase transition was observed some years ago with an attractive BEC of ^{39}K atoms, where the double-well potential was created by intersecting two pairs of laser beams [31]. Our double-well system is slightly different because the particles tunnel from one well to the other well following two different curved paths; moreover, our ellipsoidal configuration offers also the possibility of having persistent currents. For a sufficiently attractive BEC the transverse width σ of the BEC becomes smaller than l_{\perp} and there is the collapse of the single-peak configuration. For $\epsilon = 0$ the 1D NPSE predicts the collapse of this single-peak configuration at $\gamma_c = (4/3)(a/l_{\perp})$ [28, 30]. Thus, our numerical results of Fig. 3, obtained from the 1D GPE, are fully reliable under the condition $a/l_{\perp} \gg 1$. This is again the condition of a tight transverse confinement.

6 Superfluid fraction

Let us consider the effect of a boost velocity v_B on the BEC moving along the ellipse. In this case Eq. (19) is modified as follows

$$\mu \Phi = \left[\frac{1}{2m} (-i\hbar \partial_s - m v_B)^2 - \frac{\hbar^2 \kappa^2(s)}{8m} + \frac{2a_s(N-1)\hbar^2}{ml_{\perp}^2} |\Phi|^2 \right] \Phi, \quad (20)$$

Now we set

$$\Phi(s) = \frac{n(s)^{1/2}}{\sqrt{N}} e^{i\theta(s)}, \quad (21)$$

where $n(s) = N\rho(s)$ is the local number density of the BEC. Moreover, we introduce the local velocity field

$$v(s) = \frac{\hbar}{m} \partial_s \theta(s). \quad (22)$$

Inserting these formulas into Eq. (20) we obtain 1D stationary equations of zero-temperature superfluid hydrodynamics

$$\mu = -\frac{\hbar^2}{2m\sqrt{n}} \partial_s^2 \sqrt{n} + \frac{m}{2} (v - v_B)^2 - \frac{\hbar^2 \kappa(s)^2}{8m} + \frac{2a_s(1 - 1/N)\hbar^2 n}{ml_{\perp}^2}, \quad (23)$$

and also

$$\partial_s [n(v - v_B)] = 0. \quad (24)$$

Eq. (24) implies that $n(v - v_B) = J$, where J is a constant current density. This result is very interesting because it says that if $n(s)$ has spatial variations then also $v(x)$ must have spatial variations.

Inspired by Ref. [27], we now introduce the average value of the velocity $v(x)$ in a region $[a, b]$ of the ellipse as

$$\bar{v} = \frac{1}{(b-a)} \int_a^b v(s) ds, \quad (25)$$

Then, from the previous equations we obtain

$$\bar{v} = \frac{1}{(b-a)} \int_a^b \left(\frac{J}{n(s)} + v_B \right) ds = \frac{J}{\bar{n}_s} + v_B, \quad (26)$$

where

$$\bar{n}_s = \frac{1}{\frac{1}{(b-a)} \int_a^b \frac{1}{n(s)} ds}. \quad (27)$$

The number density \bar{n}_s can be interpreted as the superfluid number density of the stationary state in the spatial region $[a, b]$. Indeed, Eq. (27) is the 1D version of the formula obtained by Leggett [24] for a supersolid with spatial periodicity $(b - a)$, and recently discussed by others [25–27]. If the stationary state $\Psi(s)$ moves with the average velocity \bar{v} , its current density reads $J = \bar{n}_s (\bar{v} - v_B)$, where \bar{v} is the average velocity in the region $[a, b]$ and \bar{n}_s the corresponding superfluid number density. We can also introduce

$$\bar{n} = \frac{1}{(b-a)} \int_a^b n(s) ds \quad (28)$$

that is the average number density in the region $[a, b]$. Consequently, the superfluid fraction of the BEC in the region $[0, L]$ reads

$$f_s = \frac{\bar{n}_s}{\bar{n}} = \frac{1}{\frac{N}{L^2} \int_0^L \frac{1}{n(s)} ds}, \quad (29)$$

where $N = \int_0^L n(s) dx = L\bar{n}$. This formula can be also obtained as the response of the linear momentum of the BEC to the boost velocity \bar{v}_B , that is the non-classical translational inertia of the system [25].

In Fig. 4 we plot our numerical results of the superfluid fraction f_s as a function of the adimensional interaction strength γ for different values of the eccentricity ϵ of the ellipse. For positive values of γ the superfluid fraction f_s is close to 1 also with $\epsilon = 0.9$. However, for $0 \leq \gamma < 1$ and a very large eccentricity ($\epsilon = 0.99$) we find $f_s \simeq 0.95$. This result is quite reasonable because the wavefunction is strongly localized in the two well of the quantum-curvature potential. For negative values of γ the most interesting effect appears in Fig. 4:

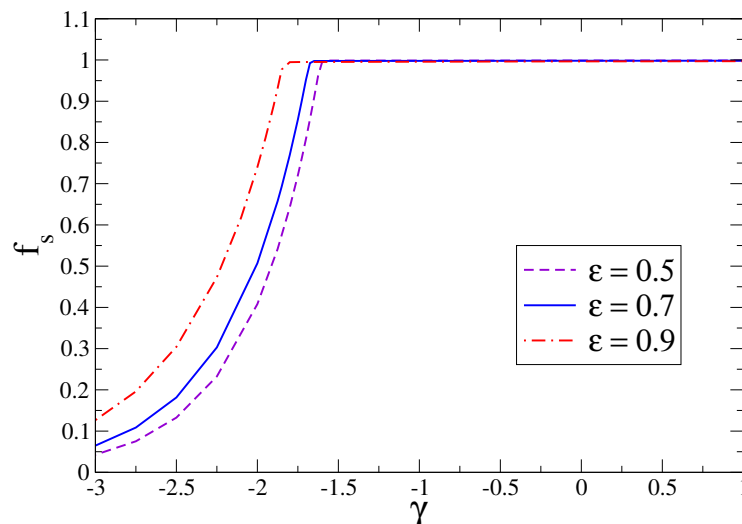


Figure 4: Superfluid fraction f_s of the BEC ground state in an ellipse as a function of the adimensional interaction strength $\gamma = 2aa_s(N - 1)/l_\perp^2$, where a is the length of the major semi-axis, a_s is the 3D s -wave scattering length and l_\perp is the characteristic length of the transverse harmonic confinement. The curves are obtained for three values of the eccentricity ϵ of the ellipse.

around $\gamma \simeq -1.6$ the superfluid fraction f_s quickly decreases and it goes to zero for very large negative values of γ . This is exactly the quantum phase transition from a two-peak configuration to a one-peak configuration. As previously stressed, the one-peak configuration becomes modulationally unstable when at least one of the energies of its elementary excitations acquires an imaginary component. A similar modulational instability [20] happens in the formation of a train of bright solitons from a single-peak Bose-Einstein condensate, induced by a sudden change in the sign of the scattering length from positive to negative [32]. Our Fig. 4 reveals that the critical strength γ_c crucially depends on the eccentricity ϵ , such as the behavior of f_s as a function of γ for $\gamma < \gamma_c$. In particular, we find that γ_c slightly reduces by increasing ϵ , but this effect is quite weak.

An important remark is that Eq. (29) has been derived here without any assumption about the sign of γ . Moreover, the absence of superfluidity for $\gamma = 0$ is true only in the thermodynamic limit. In a ring there is a finite energy gap between the ground state and the first excited state also for $\gamma = 0$. The Bose system we are considering has a finite size because it is confined in a finite elliptical ring. Finally, the 3D version of Eq. (27) was proposed historically by Leggett to characterize the superfluid density of a supersolid [24], while here Eq. (27) is used to determine the superfluid density of a Bose-Einstein condensate which is not supersolid but it is instead spatially modulated due to the crucial interplay between elliptical confinement and attractive interaction.

7 Conclusions

The main goal of this paper was to understand the role of a locally-varying curvature for a Bose-Einstein condensate confined in an elliptical waveguide. The proposed setup, and the double-well quantum-curvature potential that we have found, can be experimentally achieved by using ultracold atoms, which are a paradigmatic physical platform due to the high experimental tunability of inter-atomic interactions and trapping potentials. For instance, one can trap $N = 10^4$ ultracold Rb atoms by using a rapidly moving laser beam which creates a time-averaged elliptic-shaped toroidal optical dipole potential [33]. The length a of the major semi-axis of the ellipse can be $a \simeq 100$ microns and the transverse length $l_{\perp} = 5$ microns. The scattering length a_s could be then tuned by using an external magnetic field, which induces a Fano-Feshbach resonance [34]. Despite the fact that we focused on space curvature rather than space-time curvature, we believe that our results can be of interest not only to atomic and condensed matter physics researchers, but also to a large community working on general relativity and relativistic quantum field theory.

Acknowledgements

The author thanks Francesco Ancilotto, Koichiro Furutani, Francesco Minardi, and Andrea Tononi for useful discussions.

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