

A note on the Noether-Wald and generalized Komar charges

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Abstract

We propose a simple algorithm to compute the generalized Komar charge of any exactly gauge- and diffeomorphism-invariant theory.



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Conserved charges can be used to characterize the states and the evolution of physical systems. In this short note we describe how to compute the generalized Komar charge of any exactly gauge- and diffeomorphism-invariant theory. Our goal is to provide a simple algorithm which can be easily applied to spacetime configurations with Killing isometries (cf. Eq. (21)). This is particularly useful to determine efficiently fundamental identities in black hole thermodynamics and to prove non-existence theorems for gravitational solitons and boson stars (see for instance Refs. [1–12]).

In d -dimensional General Relativity, for each of the Killing vectors k of the metric of a given vacuum solution one can construct a $(d-2)$ -form charge $K[k]$, the so-called *Komar charge* [13], which is closed on-shell¹

$$dK[k] \doteq 0. \quad (1)$$

In asymptotically-flat spacetimes, the integral of this charge at spatial infinity (S_{∞}^{d-2}) gives, up to normalization, the value of the conserved charge of the spacetime corresponding to the Killing vector: total mass/energy if k is a Killing vector that generates time translations, etc. In General Relativity, the Komar charge coincides with the Noether–Wald charge associated with the invariance under diffeomorphisms generated by vector fields ξ , $Q[\xi]$, evaluated over the Killing vector k , $Q[k]$. This fact suggests that, in more general theories (with matter or with terms of higher order in the curvature) the Komar charge may also be given by the Noether–Wald charge evaluated on k , $Q[k]$, since $Q[\xi]$ can be constructed in any theory invariant under diffeomorphisms. This naive expectation turns out to be false in general. In order to understand why and what has to be done to construct an on-shell-closed 2-form (a *generalized Komar charge*), it is convenient to review the algorithm that leads to $Q[\xi]$.

Let us consider a theory of gravity, described by the Vierbein e^a , coupled to a number of matter fields denoted generically by φ , whose dynamics is dictated by the action $S[e, \varphi]$.

¹We indicate with $\doteq$ those identities which are only satisfied on-shell.

Under a generic infinitesimal variation of the fields

$$\delta S[e, \varphi] = \int \{ \mathbf{E}_a \wedge \delta e^a + \mathbf{E}_\varphi \wedge \delta \varphi + d\Theta(e, \varphi, \delta e, \delta \varphi) \}, \quad (2)$$

where, by definition, $\mathbf{E}_a$ are the Einstein equations, $\mathbf{E}_\varphi$ are the equations of motion of the matter fields and $\Theta(e, \varphi, \delta e, \delta \varphi)$ is the *symplectic prepotential*. If the theory is exactly invariant under diffeomorphisms and any other kind of gauge transformations,²

$$\delta_\xi S[e, \varphi] = - \int d\iota_\xi \mathbf{L}, \quad (3)$$

where $\iota_\xi \mathbf{L}$ indicates the interior product of the vector field ξ with the d -form $\mathbf{L}$.³ On the other hand, if we use, instead, the general infinitesimal variation Eq. (2) particularized for infinitesimal diffeomorphisms, we get

$$\delta_\xi S[e, \varphi] = \int \{ \mathbf{E}_a \wedge \delta_\xi e^a + \mathbf{E}_\varphi \wedge \delta_\xi \varphi + d\Theta(e, \varphi, \delta_\xi e, \delta_\xi \varphi) \}. \quad (4)$$

Combining equations (3) and (4) we arrive at the identity

$$\int \{ \mathbf{E}_a \wedge \delta_\xi e^a + \mathbf{E}_\varphi \wedge \delta_\xi \varphi + d[\Theta(e, \varphi, \delta_\xi e, \delta_\xi \varphi) + \iota_\xi \mathbf{L}] \} = 0, \quad (5)$$

which is satisfied off-shell for any ξ and, more importantly, independently of the integration domain, which leads to

$$\mathbf{E}_a \wedge \delta_\xi e^a + \mathbf{E}_\varphi \wedge \delta_\xi \varphi = d[-\Theta(e, \varphi, \delta_\xi e, \delta_\xi \varphi) - \iota_\xi \mathbf{L}]. \quad (6)$$

Equation (6) tells us that the l.h.s. is a total derivative. Such total derivative can be independently evaluated integrating by parts the l.h.s. of (6). One must integrate all the terms involving derivatives of the gauge parameters until obtaining an expression that contains terms linear in the gauge parameters (without their derivatives) plus a total derivative

$$\mathbf{E}_a \wedge \delta_\xi e^a + \mathbf{E}_\varphi \wedge \delta_\xi \varphi = \mathcal{N}_a \xi^a + dN. \quad (7)$$

Comparing (6) and (7) we conclude that the terms linear in the gauge parameters must vanish⁴ and only the total derivative dN survives

$$\mathbf{E}_a \wedge \delta_\xi e^a + \mathbf{E}_\varphi \wedge \delta_\xi \varphi = dN, \quad (8)$$

where N is a $(d-1)$ -form that vanishes on-shell. Then, we can write

$$\delta_\xi S[e, \varphi] = \int d\Theta', \quad (9a)$$

$$\Theta' \equiv \Theta(e, \varphi, \delta_\xi e, \delta_\xi \varphi) + N, \quad (9b)$$

and, comparing again this expression with Eq. (3) one concludes that the $(d-1)$ -form

$$\mathbf{J}[\xi] \equiv \Theta' + \iota_\xi \mathbf{L}, \quad (10)$$

²We exclude Chern–Simons and any other terms which are invariant up to total derivatives we do not want to deal with. We must take into account that isometries induce gauge transformations, as explained in Refs. [14–16].

³We are using differential-form notation. the rest of our conventions and notation can be found in Refs. [14, 17].

⁴Notice that $\mathcal{N}_a = 0$ are nothing but the ξ -independent Noether identities. Our derivation is basically the derivation of the second Noether theorem.

is closed off-shell for any vector field ξ

$$dJ[\xi] = 0. \quad (11)$$

This implies the local existence of Noether–Wald $(d-2)$ -form $Q[k]$ defined, up to total derivatives, by

$$dQ[\xi] = J[\xi]. \quad (12)$$

Let us examine the right-hand side of this equation using the definition of $J[\xi]$, Eqs. (10), and that of Θ' , Eq. (9b)

$$dQ[\xi] = \Theta(e, \varphi, \delta_\xi e, \delta_\xi \varphi) + N + \iota_\xi L. \quad (13)$$

The second term in the right-hand side vanishes on-shell while the first vanishes if we find Killing (or reducibility [18]) parameters⁵ k such that

$$\delta_k e^a = \delta_k \varphi = 0, \quad (14)$$

because the symplectic prepotential is linear on $\delta_\xi e^a$ and $\delta_\xi \varphi$. The third term only vanishes in some simple theories like General Relativity with no matter or with free, massless scalars, and, in general, we have to deal with the equation

$$d(\mathcal{O}_s Q[k]) \doteq \mathcal{O}_s \iota_k L, \quad (15)$$

where we have introduced the on-shell-setting operator $\mathcal{O}_s$ that evaluates the expression on its right over the solution s . Smarr formulas can be computed integrating this identity over hypersurfaces with boundary on the horizon and spatial infinity [19–22]. In Refs. [1, 2] it was argued that the right-hand side of Eq. (15) is always a total derivative

$$\mathcal{O}_s \iota_k L \equiv d\omega_k, \quad (16)$$

and we can define the *generalized Komar $(d-2)$ -form charge*

$$K[k] \equiv -\mathcal{O}_s Q[k] + \omega_k, \quad (17)$$

which is closed on-shell

$$dK[k] \doteq 0. \quad (18)$$

At first sight, this construction may give a trivial $K[k]$, but the explicit calculations performed in Refs. [1–12] proof otherwise. However, a closer inspection of the way in which those explicit calculations have been performed shows that what was computed in those references is, actually, and more precisely,

$$\iota_k \mathcal{O}_s L \equiv d\omega_k, \quad (19)$$

where $\mathcal{O}_s L$ is typically obtained from the trace of the Einstein equations. The difference between this definition of ω_k and the former is a total derivative that vanished on-shell

$$\iota_k \mathcal{O}_s L - \mathcal{O}_s \iota_k L = d(\omega_k - Q[k]) \doteq 0, \quad (20)$$

and it is, precisely, the total derivative of the generalized Komar charge. Thus, we arrive at the following prescription:

$$dK[k] = [\iota_k, \mathcal{O}_s] L. \quad (21)$$

⁵In general, the transformations δ_ξ depend on the vector field ξ and other gauge parameters.

Let us see how the prescription works in the simple example of the Einstein–Maxwell theory in d dimensions. Its action is

$$S[e^a, A] = \frac{(-1)^{d-1}}{16\pi G_N^{(d)}} \int [\star(e^a \wedge e^b) \wedge R_{ab} - \frac{1}{2} F \wedge \star F] \equiv \int \mathbf{L}, \quad (22)$$

and its equations of motion and symplectic prepotential, defined by

$$\delta S = \int \{ \mathbf{E}_a \wedge \delta e^a + \mathbf{E} \wedge \delta A + d\Theta(e, A, \delta e, \delta A) \}, \quad (23)$$

are given by

$$\mathbf{E}_a = \iota_a \star (e^c \wedge e^d) \wedge R_{cd} + \frac{1}{2} (\iota_a F \wedge \star F - F \wedge \iota_a \star F), \quad (24a)$$

$$\mathbf{E} = -d \star F, \quad (24b)$$

$$\Theta(e, A, \delta e, \delta A) = -\star(e^a \wedge e^b) \wedge \delta \omega_{ab} + \star F \wedge \delta A, \quad (24c)$$

where ι_c stands for ι_{e_c} , where $e_c = e_c^\mu \partial_\mu$ and where we are ignoring the factor $(16\pi G_N^{(d)})^{-1}$ for the moment in order to get simpler expressions. A straightforward calculation using the explicit expression of the Einstein equation Eq. (24a) gives

$$\begin{aligned} (-1)^{d-1} \iota_k \mathbf{L} &= \iota_k \star (e^a \wedge e^b) \wedge R_{ab} + (-1)^d \star (e^a \wedge e^b) \wedge \iota_k R_{ab} - \frac{1}{2} \iota_k F \wedge \star F - \frac{1}{2} F \wedge \iota_k \star F \\ &= k^a \mathbf{E}_a - \iota_k F \wedge \star F + (-1)^{d-1} \star (e^a \wedge e^b) \wedge \iota_k R_{ab}. \end{aligned} \quad (25)$$

The assumption of invariance under the diffeomorphism generated by k implies the existence of the Maxwell and Lorentz momentum maps P_k and P_k^{ab} satisfying the momentum map equations

$$\iota_k R^{ab} = -\mathcal{D}P_k^{ab}, \quad (26a)$$

$$\iota_k F = -dP_k. \quad (26b)$$

The first equation is always solved by the *Killing bivector* $\mathcal{D}^a k^b$. Using these equations and integrating by parts, we find

$$\begin{aligned} (-1)^{d-1} \iota_k \mathbf{L} &= k^a \mathbf{E}_a + dP_k \wedge \star F + (-1)^{d-1} \star (e^a \wedge e^b) \wedge \mathcal{D}P_{kab} \\ &= d[-\star(e^a \wedge e^b)P_{kab} + P_k \star F] + P_k \mathbf{E} + k^a \mathbf{E}_a, \end{aligned} \quad (27)$$

so that

$$\mathcal{O}_k \iota_k \mathbf{L} \doteq d(-1)^d [\star(e^a \wedge e^b)P_{kab} - P_k \star F] = (\mathcal{O}_k \mathbf{Q}[k]). \quad (28)$$

On the other hand, taking the trace of the Einstein equation Eq. (24a)

$$\begin{aligned} e^a \wedge \mathbf{E}_a &= (d-2) \star(e^c \wedge e^d) \wedge R_{cd} - \frac{(d-4)}{2} F \wedge \star F \\ &= (d-2) [(-1)^{d-1} \mathbf{L} + \frac{1}{2} F \wedge \star F] - \frac{(d-4)}{2} F \wedge \star F \\ &= (d-2) (-1)^{d-1} \mathbf{L} + F \wedge \star F, \end{aligned} \quad (29)$$

so

$$\mathbf{L} = \frac{(-1)^d}{d-2} F \wedge \star F + \frac{(-1)^d}{d-2} e^a \wedge \mathbf{E}_a. \quad (30)$$

Then,

$$\begin{aligned} \iota_k \mathcal{O}_s \mathbf{L} &= \iota_k \left[\frac{(-1)^d}{d-2} F \wedge \star F \right] \\ &= \frac{(-1)^d}{d-2} [\iota_k F \wedge \star F + F \wedge \iota_k \star F]. \end{aligned} \quad (31)$$

We can use Eq. (26b) in the first term. In the second, profiting from the fact that we are working on-shell, we can use the dual Maxwell momentum-map equation⁶

$$\iota_k \star F \doteq -d\tilde{P}_k, \quad (32)$$

and, integrating by parts and using the equations of motion and Bianchi identities, we have

$$\begin{aligned} \iota_k \mathcal{O}_s \mathbf{L} &= \frac{(-1)^{d-1}}{d-2} [dP_k \wedge \star F + F \wedge d\tilde{P}_k] \\ &= d \left\{ \frac{(-1)^{d-1}}{d-2} [P_k \star F + \tilde{P}_k F] \right\}. \end{aligned} \quad (33)$$

Now

$$[\iota_k, \mathcal{O}_s] \mathbf{L} = d(-1)^{d-1} \left\{ \star(e^a \wedge e^b) P_{kab} - \frac{(d-3)}{d-2} P_k \star F + \frac{1}{d-2} \tilde{P}_k F \right\}, \quad (34)$$

which coincides with the result found in Ref. [4].

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References

- [1] T. Ortín, *Komar integrals for theories of higher order in the Riemann curvature and black-hole chemistry*, J. High Energy Phys. **08**, 023 (2021), doi:[10.1007/JHEP08\(2021\)023](https://doi.org/10.1007/JHEP08(2021)023) [preprint doi:[10.48550/arXiv.2107.05467](https://doi.org/10.48550/arXiv.2107.05467)].
- [2] D. Mitsios, T. Ortín and D. Pereñíguez, *Komar integral and Smarr formula for axion-dilaton black holes versus S duality*, J. High Energy Phys. **08**, 019 (2021), doi:[10.1007/JHEP08\(2021\)019](https://doi.org/10.1007/JHEP08(2021)019) [preprint doi:[10.48550/arXiv.2107.05467](https://doi.org/10.48550/arXiv.2107.05467)].
- [3] P. Meessen, D. Mitsios and T. Ortín, *Black hole chemistry, the cosmological constant and the embedding tensor*, J. High Energy Phys. **12**, 155 (2022), doi:[10.1007/JHEP12\(2022\)155](https://doi.org/10.1007/JHEP12(2022)155) [preprint doi:[10.48550/arXiv.2211.12231](https://doi.org/10.48550/arXiv.2211.12231)].
- [4] T. Ortín and D. Pereñíguez, *Magnetic charges and Wald entropy*, J. High Energy Phys. **11**, 081 (2022), doi:[10.1007/JHEP11\(2022\)081](https://doi.org/10.1007/JHEP11(2022)081) [preprint doi:[10.48550/arXiv.2211.12231](https://doi.org/10.48550/arXiv.2211.12231)].
- [5] R. Ballesteros, C. Gómez-Fayrén, T. Ortín and M. Zatti, *On scalar charges and black hole thermodynamics*, J. High Energy Phys. **05**, 158 (2023), doi:[10.1007/JHEP05\(2023\)158](https://doi.org/10.1007/JHEP05(2023)158) [preprint doi:[10.48550/arXiv.2302.11630](https://doi.org/10.48550/arXiv.2302.11630)].
- [6] C. Gómez-Fayrén, P. Meessen, T. Ortín and M. Zatti, *Wald entropy in Kaluza-Klein black holes*, J. High Energy Phys. **08**, 039 (2023), doi:[10.1007/JHEP08\(2023\)039](https://doi.org/10.1007/JHEP08(2023)039) [preprint doi:[10.48550/arXiv.2305.01742](https://doi.org/10.48550/arXiv.2305.01742)].

⁶Notice that, in d dimensions, $\tilde{P}_k$ is a $(d-4)$ -form.

- [7] I. Bandos and T. Ortín, *Noether-Wald charge in supergravity: The fermionic contribution*, J. High Energy Phys. **12**, 095 (2023), doi:[10.1007/JHEP12\(2023\)095](https://doi.org/10.1007/JHEP12(2023)095) [preprint doi:[10.48550/arXiv.2305.10617](https://doi.org/10.48550/arXiv.2305.10617)].
- [8] R. Ballesteros and T. Ortín, *Hairy black holes, scalar charges and extended thermodynamics*, Class. Quantum Gravity **41**, 055007 (2024), doi:[10.1088/1361-6382/ad210a](https://doi.org/10.1088/1361-6382/ad210a).
- [9] R. Ballesteros and T. Ortín, *Generalized Komar charges and Smarr formulas for black holes and boson stars*, SciPost Phys. Core **8**, 038 (2025), doi:[10.21468/SciPostPhysCore.8.2.038](https://doi.org/10.21468/SciPostPhysCore.8.2.038) [preprint scipost.org/submissions/scipost_202411_00021v2/].
- [10] I. Bandos, P. Meessen and T. Ortín, *Noether-Wald and Komar charges in supergravity, fermions, and Killing supervectors in superspace*, J. Phys.: Conf. Ser. **2912**, 012007 (2024), doi:[10.1088/1742-6596/2912/1/012007](https://doi.org/10.1088/1742-6596/2912/1/012007).
- [11] T. Ortín and M. Zatti, *On the thermodynamics of the black holes of the Cano-Ruipérez 4-dimensional string effective action*, J. High Energy Phys. **06**, 026 (2025), doi:[10.1007/JHEP06\(2025\)026](https://doi.org/10.1007/JHEP06(2025)026) [preprint doi:[10.48550/arXiv.2411.10417](https://doi.org/10.48550/arXiv.2411.10417)].
- [12] I. Bandos, P. Meessen and T. Ortín, *Komar charge of $\mathcal{N} = 2$ supergravity and its superspace generalization*, J. High Energy Phys. **05**, 231 (2025), doi:[10.1007/JHEP05\(2025\)231](https://doi.org/10.1007/JHEP05(2025)231) [preprint doi:[10.48550/arXiv.2412.18510](https://doi.org/10.48550/arXiv.2412.18510)].
- [13] A. Komar, *Covariant conservation laws in general relativity*, Phys. Rev. **113**, 934 (1959), doi:[10.1103/PhysRev.113.934](https://doi.org/10.1103/PhysRev.113.934).
- [14] Z. Elgood, P. Meessen and T. Ortín, *The first law of black hole mechanics in the Einstein-Maxwell theory revisited*, J. High Energy Phys. **09**, 026 (2020), doi:[10.1007/JHEP09\(2020\)026](https://doi.org/10.1007/JHEP09(2020)026) [preprint doi:[10.48550/arXiv.2006.02792](https://doi.org/10.48550/arXiv.2006.02792)].
- [15] Z. Elgood, D. Mitsios, T. Ortín and D. Pereñíguez, *The first law of heterotic stringy black hole mechanics at zeroth order in α'* , J. High Energy Phys. **07**, 007 (2021), doi:[10.1007/JHEP07\(2021\)007](https://doi.org/10.1007/JHEP07(2021)007).
- [16] Z. Elgood, T. Ortín and D. Pereñíguez, *The first law and Wald entropy formula of heterotic stringy black holes at first order in α'* , J. High Energy Phys. **05**, 110 (2021), doi:[10.1007/JHEP05\(2021\)110](https://doi.org/10.1007/JHEP05(2021)110) [preprint doi:[10.48550/arXiv.2012.14892](https://doi.org/10.48550/arXiv.2012.14892)].
- [17] T. Ortín, *Gravity and strings*, Cambridge University Press, Cambridge, UK, ISBN 9780521768139 (2015), doi:[10.1017/CBO9781139019750](https://doi.org/10.1017/CBO9781139019750).
- [18] G. Barnich and F. Brandt, *Covariant theory of asymptotic symmetries, conservation laws and central charges*, Nucl. Phys. B **633**, 3 (2002), doi:[10.1016/S0550-3213\(02\)00251-1](https://doi.org/10.1016/S0550-3213(02)00251-1) [preprint doi:[10.48550/arXiv.hep-th/0111246](https://doi.org/10.48550/arXiv.hep-th/0111246)].
- [19] D. Kastor, *Komar integrals in higher (and lower) derivative gravity*, Class. Quantum Gravity **25**, 175007 (2008), doi:[10.1088/0264-9381/25/17/175007](https://doi.org/10.1088/0264-9381/25/17/175007) [preprint doi:[10.48550/arXiv.0804.1832](https://doi.org/10.48550/arXiv.0804.1832)].
- [20] D. Kastor, S. Ray and J. Traschen, *Smarr formula and an extended first law for Lovelock gravity*, Class. Quantum Gravity **27**, 235014 (2010), doi:[10.1088/0264-9381/27/23/235014](https://doi.org/10.1088/0264-9381/27/23/235014) [preprint doi:[10.48550/arXiv.1005.5053](https://doi.org/10.48550/arXiv.1005.5053)].

- [21] S. Liberati and C. Pacilio, *Smarr formula for Lovelock black holes: A Lagrangian approach*, Phys. Rev. D **93**, 084044 (2016), doi:[10.1103/PhysRevD.93.084044](https://doi.org/10.1103/PhysRevD.93.084044) [preprint doi:[10.48550/arXiv.1511.05446](https://doi.org/10.48550/arXiv.1511.05446)].
- [22] H. Adami, M. Golshani, M. M. Sheikh-Jabbari, V. Taghiloo and M. H. Vahidinia, *Covariant phase space formalism for fluctuating boundaries*, J. High Energy Phys. **09**, 157 (2024), doi:[10.1007/JHEP09\(2024\)157](https://doi.org/10.1007/JHEP09(2024)157) [preprint doi:[10.48550/arXiv.2407.03259](https://doi.org/10.48550/arXiv.2407.03259)].